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IN

He ma sm cor ger $\{f_t\}$ on

PR

If f*t

$$\mathcal{A}(t) := \mathcal{A}[g_t, \Phi_t; \mathcal{U}] = \mathcal{A}[g, \Phi; \mathcal{U}]$$

sin So hav

$$0 \; = \; \frac{d\mathcal{A}}{dt}|_{t=0} \; = \; \int_{\mathcal{U}} \; \{ \; -\frac{1}{2} \; T^{\mu\nu} \; \dot{g}_{\mu\nu} \; + \; E \; \dot{\Phi} \; \} \; d\mu_g$$

Е res

$$\dot{g} = \mathcal{L}_X g$$
 , $\dot{g}_{\mu
u} = \nabla_\mu X_
u + \nabla_
u X_\mu$

EULER-LAGRANGE EQUATIONS WITH APPLICATION TO MATTER FIELDS

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ABSTRACT:

This paper introduces the basic concept of Euker Lagrange equation. Our problem is $\nabla_{\nu}T^{\mu\nu} = 0$.for matter fields.

In this paper, step has been taken to prove it numerically.

YWORDS : collection of atter, basic concept.	Hence $0 = -\int \frac{1}{2} T^{\mu\nu} (\nabla_{\mu} X_{\nu} + \nabla_{\nu} X_{\mu}) d\mu_{q}$	Finally, the gravitational stress may be written as
TRODUCTION:- re Φ is the collection of itter fields and we consider	$J_{\mathcal{U}} 2$ $= -\int_{\mathcal{U}} T^{\mu\nu} \nabla_{\nu} X_{\mu} d\mu_{g}$ $= \int_{\mathcal{U}} (\nabla_{\nu} T^{\mu\nu}) X_{\mu} d\mu_{g}.$	$T^{\mu\nu} = 4 \frac{\partial L^*}{\partial \alpha} F^{\mu}_{\kappa} F^{\nu\kappa} + \left(\beta \frac{\partial L^*}{\partial \beta} - L^*\right) g^{\mu\nu}$
ooth vector field X with mpact support in U. This nerates a 1-parameter-group	X is arbitrary smooth of compact	We get
of diffeomorphisms of U to itself. [1-5]	So we write the gravitational stress in the following way.	$L^* = \pi^{\mu u} \pi_{\mu u}$
OOF	${\partial \sigma\over\partial g^{\mu u}}~=~\partial_{\mu}~\Phi~\partial_{ u}~\Phi$	When g vary, we can write
g and Φ by $\Phi_t = f_t^*$ Then it is	$T^{\mu u} = 2 \; {dL^*\over d\sigma} \; \partial^\mu \; \Phi \; \partial^ u \; \Phi \; - \; L^* \; g^{\mu u}$	$\pi = \mathcal{L}_U g$
$(t) := \mathcal{A} [g_t, \Phi_t; \mathcal{U}] = \mathcal{A} [g, \Phi; \mathcal{U}]$	Electromagnetic field.	$\int_{U} L^* d\mu_g$, then it is
ce the action is an invariant. A(t) is independent of t. We	First	
Ve $= \frac{dA}{dt} _{t=0} = \int_{\mathcal{U}} \{ -\frac{1}{2} T^{\mu\nu} \dot{g}_{\mu\nu} + E \dot{\Phi} \} d\mu_g =$	$\frac{\partial \alpha}{\partial g^{\mu\nu}} = 2 F_{\mu\kappa} F^{\kappa}_{\nu}$	
is the variation of A with	and $\partial \beta = \frac{1}{2} \rho_{\alpha}$ because of the identity $E = *E \delta = \frac{1}{2} \rho_{\alpha}$	
	$\frac{\partial g^{\mu\nu}}{\partial g^{\mu\nu}} = \frac{2}{2} \beta g_{\mu\nu}$, because of the identity: $F_{\mu\kappa} \cdot F_{\nu}^{\nu} = \frac{1}{4} \beta g$	μ

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$$\begin{split} \dot{\mathcal{A}} &= \int_{\mathcal{U}} \left\{ \left(-2 \ \pi_{\kappa}^{\mu} \ \pi^{\nu\kappa} + \frac{1}{2} \ \pi^{\kappa\lambda} \ \pi_{\kappa\lambda} \ g^{\mu\nu} \right) \dot{g}_{\mu\nu} \ + \ 2 \ \pi^{\mu\nu} \ \dot{\pi}_{\mu\nu} \ \right\} \ d\mu_g \ , \\ \dot{\pi}_{\mu\nu} &= \mathcal{L}_U \dot{g}_{\mu\nu} \ = \ U^{\kappa} \nabla_{\kappa} \ \dot{g}_{\mu\nu} \ + \ \dot{g}_{\kappa\nu} \ \nabla_{\mu} \ U^{\kappa} \ + \ \dot{g}_{\kappa\mu} \ \nabla_{\nu} \ U^{\kappa} \ , \\ \int_{\mathcal{U}} \ \pi^{\mu\nu} \ \dot{\pi}_{\mu\nu} \ d\mu_g \ = \ \int_{\mathcal{U}} \left\{ -\nabla_{\kappa} \ (U^{\kappa} \ \pi^{\mu\nu}) \ + \ \pi^{\kappa\nu} \ \nabla_{\kappa} \ U^{\mu} \ + \ \pi^{\kappa\mu} \ \nabla_{\kappa} \ U^{\nu} \ \right\} \dot{g}_{\mu\nu} \ d\mu_g \ . \end{split}$$

So, we get

$$T^{\mu\nu} = 4 \pi^{\mu}_{\kappa} \pi^{\nu\kappa} - \pi^{\kappa\lambda} \pi_{\kappa\lambda} g^{\mu\nu} + 4 \left\{ \nabla_{\kappa} \left(U^{\kappa} \pi^{\mu\nu} \right) - \pi^{\kappa\nu} \nabla_{\kappa} U^{\mu} - \pi^{\kappa\mu} \nabla_{\kappa} U^{\nu} \right\}$$

Now, we assume the case of conformal properties.

Let

$$g \mapsto \tilde{g} = \Omega^2 g$$
, $\mathcal{A}[\tilde{g}] = \mathcal{A}[g]$

Hence

$$\dot{g} = \lambda g , \quad \lambda = 2 \Omega \dot{\Omega}$$

And

$$\dot{\mathcal{A}} = 0$$

We know

$$\dot{\mathcal{A}} = -\frac{1}{2} \int_{\mathcal{U}} T^{\mu\nu} \dot{g}_{\mu\nu} d\mu_g$$

we get.

$$0 ~=~ -\frac{1}{2} ~\int_{\mathcal{U}} ~\lambda ~tr ~T ~d\mu_g$$

For arbitrary function λ , it follows that trT = 0.

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