# **ORIGINAL ARTICLE**





# FIVE DIMENSIONAL PLANE GRAVETATIONAL WAVES IN BIMETRIC RELATIVITY

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#### **Abstract:**

The plane gravitational wave solutions of the field equations  $N_i^j=0$  in  $V_5$  for bimetric relativity are given by  $g_{ij}$  satisfying  $Q\rho_i^j+P\sigma_i^j=0$ 

### **KEYWORDS:**

Five Dimensional, plane gravitational wave, Bimetric Relativity.

### [1] INTRODUCTION

In the paper refer it to [1], for bimetric relativity of Rosen (1973,74), he has obtained the plane wave solutions  $g_{ij}$  of the field equations  $N_i^j = 0$  in  $V_5$  by reformulating Adhav and Karade's (1994) definition of plane wave in four dimensional space-time  $V_4$  as follows:

#### **DEFINITIONS**

A plane gravitational wave  $g_{ij}$  is a non-flat solution of the field equations

$$N_i^j = 0$$
 (i,j = 1,2,3,4,5) (1.1)

In an empty region of space-time such that

$$g_{ij} = g_{ij}(Z), \qquad Z = Z(x^i), \quad x^i = u, x, y, z, t$$
 (1.2)

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in some suitable co-ordinate system such that

$$g^{ij}Z_{,i}Z_{,i} = 0, \quad Z_{,i} = \frac{\partial Z}{\partial x^i}$$
 (1.3)

$$Z = Z(z,t), \quad Z_{,4} = 0, \quad Z_{,5} = 0$$
 (1.4)

where 
$$N_i^j = \frac{1}{2} f^{\alpha\beta} (g^{hj} g_{hi/\alpha})_{/\beta}$$

$$N = N_i^j$$
,  $K = \sqrt{g/f}$ ,  $g = \det(g_{ij})$ ,  $f = \det(f_{ij})$ 

and the bar (/) stands for f-covariant differentiation.

In this definition, the signature convention adopted is

$$g_{aa} < 0$$
,  $a = 1,2,3,4$ 

$$\begin{vmatrix} g_{aa} & g_{ab} \\ g_{ba} & g_{bb} \end{vmatrix} > 0, \quad \begin{vmatrix} g_{11} & g_{12} & g_{13} \\ g_{21} & g_{22} & g_{23} \\ g_{31} & g_{32} & g_{33} \end{vmatrix} < 0$$

$$\begin{vmatrix} g_{11} & g_{12} & g_{13} & g_{14} \\ g_{21} & g_{22} & g_{23} & g_{24} \\ g_{31} & g_{32} & g_{33} & g_{34} \\ g_{41} & g_{42} & g_{43} & g_{44} \end{vmatrix} > 0, \quad g_{55} > 0,$$

$$(1.5)$$

(not summed for 
$$a, b = 1, 2, 3, 4$$
) and accordingly  $g = \det(g_{ij}) > 0.$  (1.6)

Though the definition of plane wave used in paper [1] is similar to that of Adhao and Karade (1994), the result (1.6) is contrary to that of Adhao and Karade which is obvious because of the signature difference and odd dimension of the space-time.

The field equations  $N_i^j = 0$  then yield

$$N\rho_i^j + M\sigma_i^j = 0,$$

Which is equivalent to

$$= \overline{w_1} \rho_i^j + \overline{w_1} \sigma_i^j = 0 = \overline{\phi_1} \rho_i^j + \overline{\phi_1} \sigma_i^j$$

Where

$$w = t + z\phi, \qquad \phi = \frac{Z_{,4}}{Z_{,5}}.$$

$$M = \overline{w} - z\overline{\phi}, \qquad N = \overline{w} - z\overline{\phi}$$

$$\rho_i^j = [\phi^2 - 1]g^{hj}\overline{g}_{hi}$$

$$\sigma_i^j = \frac{d}{dZ}[(1 - \phi^2)g^{hj}\overline{g}_{hi}]$$

It is observed that the format of mathematical expression derived by Adhao and Karade (1994) in  $V_4$  is retained even in five dimensional space-time  $V_5$  for BR theory of Rosen (1973,74), where at each point of the space-time there are two metrics

$$ds^2 = g_{ii}dx^i dx^i$$

and 
$$d\sigma^2 = f_{ij}dx^i dx^i$$

In the present paper, same space-time is considered as in the paper [1] but relax the conditions (1.2), (1.3) and (1.5) with assuming

$$Z = Z(y, z, t), \quad Z_{,3} \neq 0, \quad Z_{,4} \neq 0, \quad Z_{,5} \neq 0$$
 (1.7)

we get some interesting result in bimetric theory of relativity.

# [2] SOLUTIONS OF FIELD EQUATIONS

From the equations (1.3) and (1.7), we get

$$g^{33}\phi_1^2 + 2g^{34}\phi_1\phi_2 + 2g^{35}\phi_1 + g^{44}\phi_2^2 + g^{45}\phi_2 + g^{55} = 0$$
 (2.1)

Where 
$$\phi_1 = \frac{Z_{3}}{Z_{5}}, \phi_2 = \frac{Z_{4}}{Z_{5}}$$
 (2.2)

which further yield

$$t + \phi_1 y = w_1(Z) \tag{2.3}$$

$$t + \phi_2 z = w_1(Z) \tag{2.4}$$

Where  $w_1$  and  $w_2$  are an arbitrary functions of Z.

Differentiating partially (2.3) with respect to y, t and (2.4) with respect to z, t, we obtain

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$$Z_{,3} = \frac{\phi_1}{M_1}$$
,  $Z_{,5} = \frac{1}{M_1}$  (2.5)

where  $M_1 = \overline{w_1} - \overline{\phi_1} y$  (2.6)

$$Z_{4} = \frac{\phi_2}{M_2}$$
,  $Z_{5} = \frac{1}{M_2}$  (2.7)

where 
$$M_2 = \overline{w}_2 - \overline{\phi}_2 z$$
 (2.8)

Differentiating partially (2.6) with respect to y, t and (2.8) with respect to z, t, we obtain

$$M_{1'3} = \frac{N_1}{M_1} \phi_1 - \overline{\phi_1} \quad , \qquad M_{1'5} = \frac{N_1}{M_1}$$
 (2.9)

Where 
$$N_1 = w_1 - \phi_1 y$$
 (2.10)

$$M_{2\cdot 4} = \frac{N_2}{M_2} \phi_2 - \overline{\phi_2} , \quad M_{2\cdot 4} = \frac{N_2}{M_2}$$
 (2.11)

Where 
$$N_2 = \overline{w_2} - \overline{\phi_2} y$$
 (2.12)

From equations (2.5), (2.7) and (2.9), (2.11) it is observed that

$$M_1 = M_2 = P$$
 (say) and  $N_1 = N_2 = Q$  (say) (2.13)

Using equations (2.13) the equations (2.5), (2.7) and (2.9), (2.11) can be rewritten as

$$Z_{,3} = \frac{\phi_1}{P}, \quad Z_{,4} = \frac{\phi_2}{P}, \quad Z_{,5} = \frac{1}{P}$$
 (2.14)

And 
$$P_{\cdot 3} = \frac{Q}{P} \phi_1 - \overline{\phi_1}$$
,  $P_{\cdot 4} = \frac{Q}{P} \phi_2 - \overline{\phi_2}$ ,  $P_{\cdot 5} = \frac{Q}{P}$  (2.15)

Where a bar (-) over a letter means the derivative with respect to Z.

It is to be noted that the expression of various quantities obtained here retain their forms same as in paper [1].

Presuming  $f_{ij}$  as Lorentz metric (-1,-1,-1,+1), the f-covariant derivative becomes the ordinary partial derivative and the field equation(1.1) assume the simple form

$$f^{\alpha\beta}(g^{hj}g_{hi,\alpha}),_{\beta} = 0 \tag{2.16}$$

Which in view of (1.7) becomes

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$$f^{33}(g^{hj}g_{hi'3})_{,3}+f^{44}(g^{hj}g_{hi4})_{,4}+f^{55}(g^{hj}g_{hi5})_{,5}=0$$

Which further yield

$$Q\{[(\phi_{1}^{2} + \phi_{2}^{2}) - 1]g^{hj}\overline{g}_{hi}\} + P\{[1 - (\phi_{1}^{2} + \phi_{2}^{2})]g^{hj}\overline{g}_{hi} + [1 - (\phi_{1}^{2} + \phi_{2}^{2})]g^{hj}\overline{g}_{hi} + [1 - (\phi_{1}^{2} + \phi_{2}^{2})]g^{hj}\overline{g}_{hi} + [1 - (\phi_{1}^{2} + \phi_{2}^{2})]g^{hj}\overline{g}_{hi} - [2(\phi_{1}\overline{\phi}_{1} + \phi_{2}\overline{\phi}_{2})g^{hj}\overline{g}_{hi}]\} = 0$$

$$(2.17)$$

Equation (2.17) can be put in the form analogous of Adhav and Karade (1994) as

$$Q\rho_i^j + P\sigma_i^j = 0 (2.18)$$

Where

$$\rho_i^j = [(\phi_1^2 + \phi_2^2) - 1]g^{hj}\overline{g}_{hi}$$

$$\sigma_i^j = \frac{d}{dZ} \{ [1 - (\phi_1^2 + \phi_2^2)] g^{hj} g_{hi}^{-1} \}$$

Substituting the values of P and Q, equation (2.18) reduces to

$$\overline{\overline{w}_{1}}\rho_{i}^{j} + \overline{\overline{w}_{1}}\sigma_{i}^{j} = 0 = \overline{\overline{\phi}_{1}}\rho_{i}^{j} + \overline{\phi_{1}}\sigma_{i}^{j}$$

$$\overline{\overline{w}_{2}}\rho_{i}^{j} + \overline{\overline{w}_{2}}\sigma_{i}^{j} = 0 = \overline{\phi_{2}}\rho_{i}^{j} + \overline{\phi_{2}}\sigma_{i}^{j}$$
(2.19)

Which are again in the corresponding format of Adhav and Karade (1994). However, instead of single equation obtained in the paper [1], we have two equations (2.19) due to postulation of  $Z_{,3} \neq 0$ ,  $Z_{,4} \neq 0$ ,  $Z_{,5} \neq 0$ 

# **CONCLUSION**

We conclude that the plane gravitational waves  $g_{ij}$  given by (2.18) or (2.19) are the solutions of the field equations in bimetric relativity. If Z is independent of the variable y, then the work regarding plane gravitational waves in five dimensional space – time in paper [1] can be obtained.

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